

# Rate-Splitting Multiple Access for Satellite Short-Packet Communications: Finite Blocklength Modeling and Reliability Analysis

Sang-Quang Nguyen<sup>(1)(\*)</sup>, and Chi-Bao Le<sup>(2)</sup>

**Abstract**—Short-packet transmission is becoming crucial for satellite services that cannot rely on long codewords to hit the required latency and reliability. This study investigates rate-splitting multiple access (RSMA) in that context and builds a finite-blocklength (FBL) model for a downlink satellite-terrestrial link affected by Shadowed-Rician fading. We obtain closed-form approximations for the block error rate (BLER) of both the common and private streams, explicitly incorporating imperfect successive interference cancellation (ipSIC) at the receivers. Compared with power-domain non-orthogonal multiple access (NOMA), RSMA exhibits more stable BLER the common stream helps dampen residual interference due to ipSIC and the short-packet effect—so RSMA generally needs less transmit power to attain the same error targets. Numerical results validate the analysis and demonstrate consistent RSMA advantages across a wide range of transmit powers, blocklengths, shadowing severities, and antenna configurations. The results suggest that RSMA is a very promising option for future satellite systems that need to provide reliable, low-latency, and short-packet communications in view of realistic SIC imperfections.

**Index Terms**—Rate-splitting multiple access, finite blocklength, short-packet communications, satellite-terrestrial systems, block error rate, imperfect successive interference cancellation, reliability analysis.

## I. INTRODUCTION

Satellite-terrestrial communication systems have attracted significant interest as future 6G networks begin to demand wider coverage, massive connectivity, and reliable links for devices scattered across remote or hard-to-reach regions [1]–[3]. Unlike traditional broadband satellite services that focus on large codewords and high throughput, many emerging applications are now dominated by short-packet transmission, where only a few hundred symbols are available to encode critical information. Examples include environmental monitoring sensors, safety-related signaling, UAV telemetry packets, and low-latency control messages for distributed IoT systems [4], [5]. These scenarios impose tight reliability and delay requirements, making it difficult to rely solely on classical Shannon-capacity results, which assume infinitely long blocklengths and often mask the performance degradation triggered by short packets [6]–[8]. As a result, finite-blocklength (FBL)

analysis has become increasingly important for understanding the practical behavior of satellite links operating under strict latency constraints [9].

Another challenge that satellite systems must contend with is interference management. With the growth of spectrum sharing between terrestrial and non-terrestrial networks, serving multiple ground users simultaneously requires access schemes that can sustain reliability in the presence of strong co-channel interference. Non-orthogonal multiple access (NOMA) has been widely studied for this purpose, but its dependence on accurate successive interference cancellation (SIC) often leads to performance loss when SIC becomes imperfect—something that occurs frequently in Shadowed-Rician fading, user mobility, or channel estimation uncertainty [10]–[13]. Even small SIC errors may propagate and severely deteriorate reliability, particularly when packets are short and decoding margins are tight [14], [15]. In contrast, rate-splitting multiple access (RSMA) has emerged as a more flexible strategy that can blend partial interference decoding with partial interference treating-as-noise [16]. By dividing messages into a common stream and several private streams, RSMA can distribute interference more evenly across users and reduce the system’s sensitivity to imperfect SIC. Several recent studies have shown that RSMA tends to provide more stable throughput and outage performance than NOMA in satellite or integrated satellite-terrestrial networks [17], especially when channel conditions fluctuate or users have highly unbalanced link qualities. However, most of these studies evaluate RSMA under the infinite blocklength assumption, where error probabilities naturally vanish at high SNR and decoding imperfections are less visible. This leaves an important open question unanswered: How does RSMA behave in short-packet satellite links where reliability, not capacity, becomes the primary metric?

But another layer of complexity arises from the Shadowed-Rician fading environment that characterizes many land-mobile satellite links. Driven by the degree of shadowing—from average to heavy-deep signal fluctuations make the block error rate swing wildly, even when the average signal-to-noise ratio is the same. These swings are even tougher to manage under FBL constraints, where tight margins around the coding gain come into play. Understanding the interplay of RSMA, FBL limits, imperfect SIC, and Shadowed-Rician fading is critical to shaping robust next-generation satellite-terrestrial systems. In this spirit, the present paper provides a detailed framework for the study of RSMA-enabled satellite short-packet communication under FBL constraints. The scaling of

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reliability with transmit power, codeword length, shadowing severity, and power-splitting choices is derived, as well as a comparison of RSMA against NOMA to uncover practical differences in robustness. These contributions shall serve as guidelines for the system designer to identify regimes where RSMA yields actual gains and where its common-message approach stabilizes error performance in realistic satellite environments. Different from existing RSMA finite-blocklength studies that primarily consider terrestrial or uplink systems, and RSMA-based satellite works that typically assume infinite blocklength, this paper focuses on downlink low Earth orbit (LEO) satellite communications under finite-blocklength constraints, Shadowed-Rician fading, and imperfect SIC.

A. Motivations and Contributions

The main contributions of this paper are outlined as follows:

- We develop a finite-blocklength reliability framework for downlink RSMA-enabled LEO satellite systems over Shadowed-Rician channels, which captures the decoding behavior of both common and private messages under short-packet constraints while explicitly accounting for imperfect SIC, a combination that has not been jointly investigated in existing works.
- Closed-form approximations for the BLER of RSMA streams are derived, explicitly incorporating the impact of imperfect SIC (ipSIC), which is a realistic challenge for practical receivers in satellite environments.
- A detailed comparison with NOMA is presented, showing when and why RSMA provides more reliable short-packet transmission-particularly at moderate SNRs, under severe shadowing, or when the blocklength becomes tight.
- Extensive numerical results validate the analysis and illustrate how transmit power, antenna configuration, blocklength, and shadowing severity jointly shape the reliability of RSMA-based satellite links.

Compared to existing RSMA studies under the finite-blocklength regime, which mainly focus on terrestrial or uplink scenarios, this work investigates downlink RSMA transmission in LEO satellite systems over Shadowed-Rician channels. Moreover, unlike prior satellite RSMA analyses that typically rely on infinite blocklength assumptions, we explicitly characterize finite-blocklength BLER performance while accounting for imperfect SIC effects on both common and private streams.

B. Organization & Notations

The remainder of this paper is structured as follows. Section II describes the system model, including the satellite-terrestrial channel characteristics and the RSMA transmission strategy. Section III develops the finite-blocklength analysis and derives the corresponding expressions for the achievable performance. Section IV reports and discusses the numerical results, highlighting the comparison between RSMA and conventional multiple-access schemes. Finally, Section V summarizes the main findings and outlines potential directions for future work.

**Notations:** Throughout this paper, we use several frequently used mathematical symbols. The operator  $\mathbb{E}\{\cdot\}$  denotes expectation;  $J_i(\cdot)$  refers to the Bessel function of the first kind with order  $i$ ;  $\|\cdot\|_F$  is the Frobenius norm;  ${}_1F_1(\cdot; \cdot; \cdot)$  represents the confluent hypergeometric function of the first kind;  $(\cdot)_t$  is the Pochhammer symbol;  $\mathcal{B}(\cdot, \cdot)$  corresponds to the Beta function;  $\Gamma(\cdot)$  is the Gamma function;  $\gamma(\cdot, \cdot)$  indicates the lower incomplete Gamma function; the Gaussian  $\mathcal{Q}$ -function is written as  $\mathcal{Q}(x) = \frac{1}{2\pi} \int_x^\infty e^{-t^2/2} dt$ ; and for any random variable  $X$ , its probability density and cumulative distribution functions are denoted by  $f_X(\cdot)$  and  $F_X(\cdot)$ , respectively.

II. SYSTEM MODEL AND TERRESTRIAL CHANNEL MODEL

A. System Description

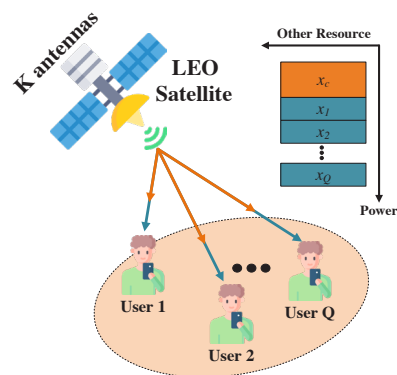


Fig. 1. An illustration of a downlink RSMA-enabled satellite-terrestrial communication system.

Fig. 1 sketches the downlink satellite-terrestrial network considered in this work, where rate-splitting multiple access (RSMA) is adopted. In this setup, a satellite  $S$  equipped with  $K$  antennas communicates simultaneously with  $Q$  ground users. The satellite operates in half-duplex mode and employs RSMA to handle interference while still aiming for better spectral efficiency. In practice, modern satellite platforms usually rely on multibeam architectures to extend coverage and capacity. For LEO systems, these beams are commonly produced by array-fed reflectors-mainly because they tend to be more efficient than direct radiating arrays. Since the beam pattern is fixed, the satellite avoids the need for heavy onboard signal processing.

The signal received at each user is shaped by the beamforming strategy, propagation environment, and power allocation used at the satellite. The channel gain from the satellite to user  $q$  is modeled as

$$\mathbf{h}_{U_q} = \mathcal{A}_q \mathbf{g}_{U_q}^\dagger \mathbf{w}_{U_q}, \tag{1}$$

where  $\mathcal{A}_q$  includes both the antenna gain and free-space path-loss. The vector  $\mathbf{g}_{U_q}$  represents the  $K \times 1$  Shadowed-Rician fading channel between the satellite antennas and user  $q$ . The operator  $(\cdot)^\dagger$  denotes the conjugate transpose, and  $\mathbf{w}_{U_q}$  is the corresponding  $K \times 1$  beamforming vector. Following the maximum ratio transmission (MRT) rule, the beamformer is

selected as  $\mathbf{w}_{U_q} = \frac{\|\mathbf{g}_{U_q}\|}{\|\mathbf{g}_{U_q}\|_F}$ , where  $\|\cdot\|_F$  is the Frobenius norm. The factor  $\mathcal{A}_q$  remains fixed and is written as

$$\mathcal{A}_q = \frac{c\sqrt{G_S G_R}}{4\pi f_c d_{S_q}}, \quad (2)$$

Here,  $c$  and  $f_c$  denoting the speed of light and the carrier frequency. The term  $d_{S_q}$  is the distance between the satellite and user  $q$ . The gains  $G_S$  and  $G_R$  correspond to the satellite beam gain and the user terminal gain, respectively. Specifically,  $G_S = G_{\max} \left[ \frac{J_1(w)}{2w} + 36 \frac{J_3(w)}{w^3} \right]^2$ , where  $G_{\max}$  is the maximum beam gain and  $w = \frac{2.07123 \sin(\theta)}{\sin(\theta_{3\text{dB}})}$ . Here,  $\theta$  denotes the angular offset from the beam center,  $\theta_{3\text{dB}}$  the 3 dB beamwidth, and  $J_1(\cdot)$ ,  $J_3(\cdot)$  are Bessel functions of order 1 and 3. In this model, different users are separated based on their channel quality obtained, for example, via pilot-assisted channel estimation. We further assume ideal CSI at the transmitter, and the results serve as a reference for future studies involving CSI imperfections [18].

### B. The signal processing at transceivers

The satellite employs RSMA to serve all users concurrently. The transmitted waveform consists of two parts: a common message, intended to be decoded by every user, and a private message for each individual user. A power coefficient  $a_c$  is assigned to the common stream, whereas the remaining power is split among the private streams.

The transmitted signal is expressed as

$$x = \sqrt{P_S} \left( \sqrt{a_c} x_c + \sum_{q=1}^Q \sqrt{a_q} x_q \right), \quad (3)$$

where  $P_S$  is the satellite transmit power,  $x_c$  is the common message with power  $a_c P_S$ , and  $x_q$  is the private message for user  $q$  with power  $a_q P_S$ . It holds that  $a_c + \sum_{q=1}^Q a_q = 1$ . The received signal at user  $q$  is then

$$\begin{aligned} y_{U_q} &= \mathbf{h}_{U_q} x + n_{U_q} \\ &= \underbrace{\mathbf{g}_{U_q}^\dagger \mathbf{w}_{U_q} \mathcal{A}_q \sqrt{a_c P_S} x_c}_{\text{Common Message}} + \underbrace{\mathbf{g}_{U_q}^\dagger \mathbf{w}_{U_q} \mathcal{A}_q \sqrt{a_q P_S} x_q}_{\text{Private Message}} \\ &\quad + \underbrace{\sum_{j=1, j \neq q}^Q \mathbf{g}_{U_q}^\dagger \mathbf{w}_{U_q} \mathcal{A}_q \sqrt{a_j P_S} x_j}_{\text{Interference}} + \underbrace{n_{U_q}}_{\text{AWGN}}, \end{aligned} \quad (4)$$

where  $n_{U_q}$  is AWGN with zero mean and variance  $N_0$ . Each user decodes the received signal in two steps.

### C. Decoding the common message

Users start by decoding the common message  $x_c$  while considering the private streams as interference. The SINR for decoding the common message at user  $q$  is

$$\begin{aligned} \bar{\gamma}_{c,q} &= \frac{a_c \mathcal{A}_q^2 P_S \|\mathbf{g}_{U_q}\|_F^2}{N_0 + \mathcal{A}_q^2 P_S (1 - a_c) \|\mathbf{g}_{U_q}\|_F^2} \\ &= \frac{a_c \mathcal{C}_q}{1 + (1 - a_c) \mathcal{C}_q}, \end{aligned} \quad (5)$$

where  $\rho_S = \frac{P_S}{N_0}$  is the signal-to-noise ratio (SNR),  $\mathcal{C}_q = \delta_q \|\mathbf{g}_{U_q}\|_F^2$  and  $\delta_q = \rho_S \mathcal{A}_q^2$ .

### D. Decoding the private message

After removing the common message, user  $q$  proceeds to decode its own private message. Interference now comes only from private messages of other users. The SINR for private decoding is

$$\bar{\gamma}_{p,q} = \frac{a_q \mathcal{C}_q}{1 + \mathcal{C}_q \sum_{j=1, j \neq q}^Q a_j + \Xi_q a_c \mathcal{C}_q}, \quad (6)$$

where  $\Xi_q$  accounts for imperfect SIC (ipSIC), with  $\Xi_q = 0$  corresponding to ideal cancellation [19].

### E. Terrestrial Channel Model

To analyze performance metrics later on, we assume the fading coefficients are i.i.d. The PDF of the channel coefficient  $g_{U_q}^{(k)}$  from the  $k$ th satellite antenna to user  $q$  is

$$f_{|g_{U_q}^{(k)}|^2}(x) = \alpha_q e^{-\beta_q x} {}_1F_1(m_q; 1; \varpi_q x), \quad x \geq 0, \quad (7)$$

where  $\alpha_q$ ,  $\beta_q$ , and  $\varpi_q$  are defined as in the original expression, with  $\Omega_q$ ,  $2b_q$ , and  $m_q$  denoting the LOS power, multipath power, and fading severity, respectively. The term  ${}_1F_1(\cdot)$  is the confluent hypergeometric function [20, Eq. (9.210.1)].

Assuming  $m_q$  takes integer values, the PDF simplifies to

$$f_{|g_{U_q}^{(k)}|^2}(x) = \alpha_q e^{-(\beta_q - \varpi_q)x} \sum_{t=0}^{m_q-1} \zeta_q(t) x^t, \quad x \geq 0, \quad (8)$$

in which  $\zeta_q(t)$  defined in terms of the Pochhammer symbol. Using [21], the PDF of  $\mathcal{C}_q$  under i.i.d. Shadowed-Rician fading is

$$f_{\mathcal{C}_q}(x) = \sum_{j_1=0}^{m_q-1} \cdots \sum_{j_K=0}^{m_q-1} \frac{\Lambda_q(K)}{\delta_q^{\Delta_q}} x^{\Delta_q-1} e^{-\left(\frac{\psi_q}{\delta_q}\right)x}, \quad (9)$$

where

$$\Lambda_q(K) = \alpha_q^K \prod_{l=1}^K \zeta_q(j_l) \prod_{u=1}^{K-1} \mathcal{B}\left(\sum_{p=1}^u j_p + u, j_{u+1} + 1\right), \quad (10)$$

and

$$\Delta_q = \sum_{l=1}^K j_l + K, \quad \psi_q = \beta_q - \delta_q. \quad (11)$$

Applying [20, Eq. (3.351.1)], the CDF of  $\mathcal{C}_q$  becomes

$$F_{\mathcal{C}_q}(x) = \sum_{j_1=0}^{m_q-1} \cdots \sum_{j_K=0}^{m_q-1} \frac{\Lambda_q(K)}{\delta_q^{\Delta_q}} \gamma\left(\Delta_q, \frac{\psi_q x}{\delta_q}\right), \quad (12)$$

which can be further simplified via [20, Eq. (8.352.6)] to

$$F_{\mathcal{C}_q}(x) = 1 - \sum_{j_1=0}^{m_q-1} \cdots \sum_{j_K=0}^{m_q-1} \sum_{p=0}^{\Delta_q-1} \frac{\Lambda_q(K) \Gamma(\Delta_q)}{p! \psi_q^{\Delta_q-p} \delta_q^p} e^{-\frac{\psi_q x}{\delta_q}} x^p. \quad (13)$$

*Remark 1:* This characterization highlights how the channel statistics influence system behavior under the Shadowed-Rician fading considered in this work.



Since the gap between  $\phi_{a,q}$  and  $\varsigma_{a,q}$  in (19) is relatively small [23],  $\tilde{\varepsilon}_{a,q}$  can be further simplified by

$$\tilde{\varepsilon}_{a,q} \approx \begin{cases} \sum_{t=1}^{\mathbb{T}} \frac{1}{\mathbb{T}} FC_q \left( \frac{\Phi_{a,q}}{a_c - \Phi_{a,q}(1-a_c)} \right), & a \in c, \\ \sum_{t=1}^{\mathbb{T}} \frac{1}{\mathbb{T}} FC_q \left( \frac{\Phi_{a,q}}{a_q - \Phi_{a,q}(\sum_{j=1, j \neq q}^Q a_j + \Xi_q a_c)} \right), & a \in p, \end{cases} \quad (22)$$

Here,  $\mathbb{T}$  implies the complexity accuracy trade-off parameter and  $\Phi_{a,q} = \phi_{a,q} + (2t-1)(\varsigma_{a,q} - \phi_{a,q})/2\mathbb{T}$ .

Using (13) and (22), we can derive the close-form of  $\tilde{\varepsilon}_{a,q}$  as

$$\tilde{\varepsilon}_{a,q} \approx 1 - \sum_{t=1}^{\mathbb{T}} \frac{1}{\mathbb{T}} \sum_{j_1=0}^{m_q-1} \cdots \sum_{j_{K-1}=0}^{m_q-1} \sum_{p=0}^{\Delta_q-1} \frac{\Lambda_q(K) \Gamma(\Delta_q)}{p! \psi_q^{\Delta_q-p} \delta_q^p} \quad (23)$$

$$\times e^{-\frac{\psi_q \xi_{a,q}}{\delta_q} \xi_{a,q}^p},$$

where  $\xi_{a,q} = \frac{\Phi_{a,q}}{a_c - \Phi_{a,q}(1-a_c)}$  for the common part ( $a = c$ ), whereas for the private part ( $a = p$ ), it is given by  $\xi_{a,q} = \frac{\Phi_{a,q}}{a_q - \Phi_{a,q}(\sum_{j=1, j \neq q}^Q a_j + \Xi_q a_c)}$ .

We derive the closed-form expression of the average BLER at each user in (17) by substituting (23) in (16).

The proof is completed.  $\blacksquare$

Reliability at  $U_q$  is the chance that the packet is accurately identified, given as [24, Eq. (27)]

$$\nu_q = (1 - \tilde{\varepsilon}_q) \times 100\%. \quad (24)$$

Furthermore, the throughput for each user may be assessed as

$$\tau_q = (1 - \tilde{\varepsilon}_{c,q}) \mathcal{R}_{c,q} + (1 - \tilde{\varepsilon}_{p,q}) \mathcal{R}_{p,q}. \quad (25)$$

#### IV. RESULTS AND DISCUSSIONS

In this segment, we conduct a series of numerical simulations to check how well the analytical expressions capture the behavior of the considered RSMA-enabled satellite-terrestrial system. The Shadowed-Rician fading parameters are chosen according to two typical shadowing conditions often used in satellite literature [25]: an average-shadowing case with  $(b_q, m_q, \Omega_q) = (0.251, 5, 0.279)$  and a heavier shadowing scenario characterized by  $(0.063, 1, 0.0007)$ . To keep the setup simple, the system involves two terrestrial users ( $Q = 2$ ),  $\mathcal{L} = \mathcal{L}_{c,q} = \mathcal{L}_{p,q} = 200$ ,  $\mathcal{T} = \mathcal{T}_{c,q} = \mathcal{T}_{p,q} = 80$  bits, both assumed to experience the same SIC imperfection level, i.e.,  $\Xi = \Xi_1 = \Xi_2$  and  $\mathbb{T} = 5$ . The equivalent noise power at each user terminal is modeled as  $N_0 = \kappa B_w T$ , where  $\kappa = 1.38 \times 10^{-23}$  denotes the Boltzmann constant,  $B_w = 50$  MHz is the transmission bandwidth, and the noise temperature is set to  $T = 290$  K, a value commonly used for LEO downlink links [26]. The satellite operates with  $K = 2$  antennas and radiates from a LEO platform toward a ground terminal located about  $d_{S_q} = 1000$  km away. The carrier frequency is  $f_c = 1.55$  GHz, and the propagation constants follow the typical free-space path-loss model. For the antenna characteristics, we adopt a receive antenna gain of  $G_R = 5$  dBi and a maximum satellite beam gain of  $G_{\max} = 25$  dBi, with an angular separation of  $\theta = 0.8^\circ$  between the users and a 3-dB beamwidth of  $\theta_{3\text{dB}} = 0.4^\circ$ . To evaluate the system

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performance under these settings, each simulation curve is averaged over  $10^6$  Monte Carlo realizations. Regarding the power distribution, the common-message power ratio is fixed at  $a_c = 0.4$ , and the remaining power is proportionally divided among the users according to  $a_1 = 0.4(1-a_c)$  and  $a_2 = 0.6(1-a_c)$ , reflecting a mild imbalance in private-message allocation. For NOMA, a fixed power allocation and the conventional SIC decoding order are assumed, which are commonly adopted for baseline comparisons.

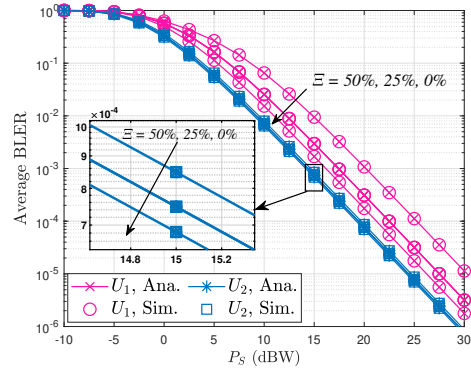


Fig. 2. Average BLER versus transmit power for different levels of SIC imperfection, with  $K = 2$ .

In Figure 2, we examine how the average BLER varies with transmit power under different levels of SIC imperfection. As expected, all curves improve gradually as the SNR increases, but the presence of residual interference still leaves a visible gap between the ideal and imperfect cases. At low SNR, the average BLER of all curves remains relatively high and close together, suggesting that noise dominates in this region. When the SNR approaches the medium range, the impact of the errors in the SIC starts becoming significant, and a small level of imprecision causes the average BLER convergence speed to reduce. For the high SNR range, it appears that the system with the highest level of imprecision for the SIC will eventually converge to the error floor, which suggests that the presence of the interference will affect the reliability of the RSMA, even when it operates at higher powers.

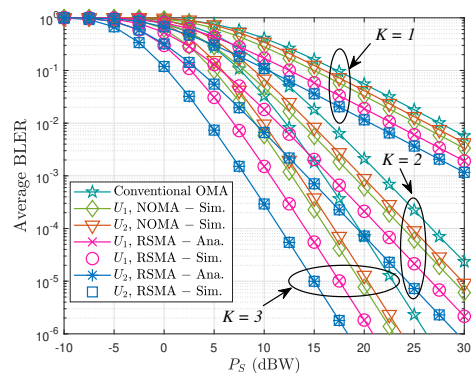


Fig. 3. Average BLER versus transmit power with increasing number of antennas at the satellite, for  $\Xi = 10\%$ .

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Figure 3 illustrates the impact of the number of satellite antennas on the BLER performance of RSMA, NOMA, and conventional OMA. As the antenna number increases, RSMA tends to achieve a more pronounced BLER reduction, benefiting more effectively from the additional array gain. In the low-SNR region, the performance of RSMA and NOMA remains close, since noise largely dominates the system behavior. Once the SNR exceeds a certain threshold, however, a clear performance separation emerges, with RSMA outperforming both NOMA and OMA. It is also apparent that  $K = 3$ , the RSMA curve drops more rapidly than those of NOMA and conventional OMA, which highlights the effectiveness of multi-antenna processing combined with rate-splitting. At higher SNR values, this performance advantage is largely preserved, while the improvement of NOMA gradually saturates due to residual interference.

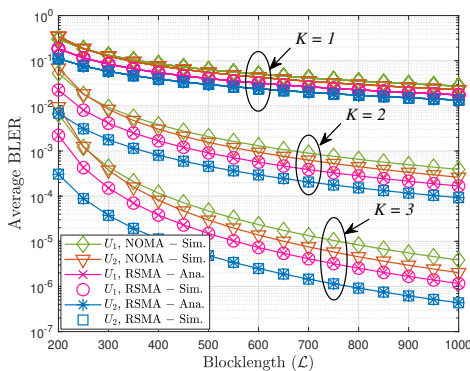


Fig. 4. Average BLER versus blocklength with increasing number of antennas at the satellite, for  $\Xi = 20\%$  and  $P_S = 10$  dBW.

Figure 4 examines the impact of blocklength on the BLER performance of RSMA and NOMA. As expected, increasing the blocklength improves reliability for both schemes due to reduced channel dispersion. However, RSMA exhibits a faster BLER decay, particularly in the short and moderate blocklength regimes, since the common message provides an additional decoding path for users experiencing unfavorable channel conditions. In contrast, NOMA degrades more rapidly when packets are short, as its performance relies heavily on reliable SIC, which becomes less effective under FBL. As the blocklength grows large, the performance gap gradually narrows and eventually saturates, indicating diminishing FBL effects. Overall, the results suggest that RSMA achieves higher reliability and better energy efficiency than NOMA in short-packet satellite transmissions.

Figure 5 illustrates the BLER behavior of the satellite link under average shadowing (AS) and heavy shadowing (HS) conditions. Under AS, where signal blockage is moderate, the channel remains relatively stable and BLER decreases rapidly with increasing transmit power. In contrast, HS introduces severe signal obstruction and deep fading, resulting in persistently higher BLER and a much slower improvement even with power boosting. These results indicate that while RSMA remains robust under AS conditions, severe shadowing im-

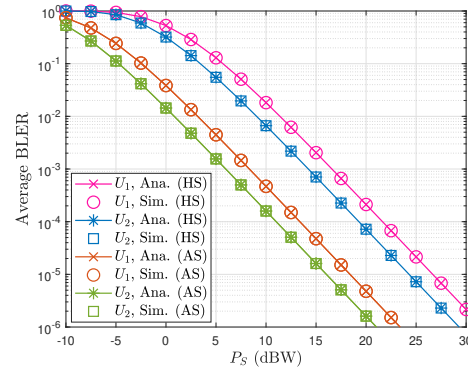


Fig. 5. Average BLER versus transmit power under various shadow fading, with  $\Xi = 10\%$  and  $K = 2$ .

poses a fundamental limitation that cannot be fully mitigated by transmit power alone.

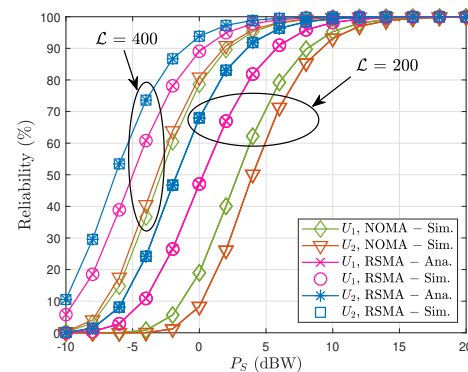


Fig. 6. Reliability versus transmit power for different blocklength values, with  $\Xi = 10\%$  and  $K = 2$ .

Figure 6 presents the reliability performance versus transmit power for different blocklengths, comparing RSMA and NOMA. For all blocklength values, increasing power enhances reliability; however, RSMA consistently achieves a target BLER at lower SNR than NOMA. This advantage becomes more pronounced for short packets, where NOMA suffers from residual interference and imperfect SIC, leading to a noticeably higher BLER in the mid-SNR region. As the blocklength increases, both schemes benefit, yet RSMA maintains a clear edge owing to its common stream, which provides additional decoding robustness. These observations confirm that RSMA is more suitable for reliability-critical short-packet satellite links.

Finally, Figure 7 compares the throughput performance of RSMA and NOMA under different satellite antenna configurations. Increasing the transmit power improves throughput for both schemes; however, RSMA consistently achieves higher throughput due to its superior decoding reliability. Moreover, the use of additional antennas enhances the SINR of both common and private streams in RSMA, allowing it to better exploit spatial diversity. At moderate SNR, RSMA exhibits a steeper throughput increase, whereas NOMA remains constrained by

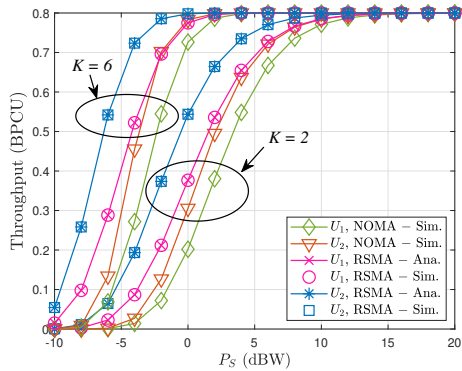


Fig. 7. Throughput versus transmit power with increasing number of antennas at the satellite, for  $\Xi = 10\%$  and  $K = \{2, 6\}$ .

residual interference and BLER effects. Even at high SNR, the throughput gap persists, highlighting that RSMA more effectively converts reliability gains into throughput benefits under short-packet conditions.

It is worth noting that the analytical and simulation results presented in this section are obtained under the assumption of perfect CSI at the receiver. In practical LEO satellite systems, channel estimation errors may arise due to mobility, Doppler effects, and limited pilot resources, which can degrade the reliability performance to some extent. Nevertheless, the results reported here serve as a useful benchmark for evaluating the fundamental behavior of RSMA under finite-blocklength transmission. Incorporating CSI imperfections into the proposed framework is an interesting extension and will be considered in future work.

### V. CONCLUSION

This paper explored how RSMA behaves in satellite short-packet communications when finite-blocklength and realistic channel conditions are taken into account. By building a tractable FBL model over Shadowed-Rician fading, we derived BLER expressions for both common and private streams, including the effect of imperfect SIC—something that often becomes a real bottleneck in practical receivers. The analysis and simulations align well and point to a consistent trend: RSMA tends to hold its reliability better than NOMA, especially when the packets are short or the shadowing becomes heavy. The common stream, although simple in structure, seems to play a meaningful role in stabilizing the error performance when the SNR is moderate and when ipSIC degrades the private layers. Numerical results further show that RSMA normally requires less transmit power to reach comparable BLER targets and scales more gracefully with blocklength and antenna count. While this study focused on a basic two-user setting, the insights suggest that RSMA is a promising direction for future satellite systems that must deliver reliable, low-latency short-packet links. Extending the model to multi-beam satellites, multi-user scenarios, or imperfect CSI would be a natural next step for future work.

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